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Local structural order in carbonic acid polymorphs: Raman and FT-IR spectroscopy

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Two different polymorphs of carbonic acid, α - and β -H₂CO₃, were identified and characterized using infrared spectroscopy (FT-IR) previously. Our attempts to determine the crystal structures of these two polymorphs using powder and thin-film X-ray diffraction techniques have failed so far. Here, we report the Raman spectrum of the α -polymorph, compare it with its FT-IR spectrum and present band assignments in line with our work on the β -polymorph [Angew. Chem. Int. Ed. 48 (2009) 2690-2694]. The Raman spectra also contain information in the wavenumber range ~90-400 cm⁻¹, which was not accessible by FT-IR spectroscopy in the previous work. While the α -polymorph shows Raman and IR bands at similar positions over the whole accessible range, the rule of mutual exclusion is obeyed for the β -polymorph. This suggests that there is a center of inversion in the basic building block of β -H₂CO₃ whereas there is none in α -H₂CO₃. Thus, as the basic motif in the crystal structure we suggest the cyclic carbonic acid dimer containing a center of inversion in case of β -H₂CO₃ and a catemer chain or a sheet-like structure based on carbonic acid dimers not containing a center of inversion in case of α -H₂CO₃. This hypothesis is strengthened when comparing Raman active lattice modes at <400 cm⁻¹ with the calculated Raman spectra for different dimers. In particular, the intense band at 192 cm⁻¹ in β -H₂CO₃ can be explained by the inter-dimer stretching mode of the centrosymmetric RC(OHO)₂ CR entity with R=OH. The same entity can be found in gas-phase formic acid (R=H) and in β -oxalic acid (R=COOH) and produces an intense Raman active band at a very similar wavenumber. The absence of this band in α -H₂CO₃ confirms that the difference to β -H₂CO₃ is found in the local coordination environment and/or monomer conformation rather than on the long range. Copyright © 2011 John Wiley & Sons, Ltd.

Keywords: carbonic acid; polymorphism; center of inversion; rule of mutual exclusion; crystal structure building blocks

Introduction

Carbonic acid and its acid-base system $CO_2/H_2O \rightleftharpoons H_2CO_3 \rightleftharpoons$ HCO_3^{-}/H^+ are of fundamental importance in many fields^[1] including astrophysics, ^[2-15] medicine, ^[16-19] marine chemistry^[20-22] and geochemistry.^[23,24] Synthesis and characterization of the intermediate carbonic acid (H₂CO₃) was long believed to be impossible owing to the two decay channels to carbon dioxide and bicarbonate. In spite of that, different routes to solid and gaseous carbonic acid have been established in the last two decades.^[1] Also, in aqueous solutions, trace amounts of H₂CO₃ were detected and characterized using spectroscopic techniques.^[25-27] Gas-phase carbonic acid was detected following the thermolysis of NH₄HCO₃ by means of mass spectrometry,^[28] following the application of a high-voltage pulse to wet CO₂ gas and supersonic jet expansion by means of FT-microwave spectroscopy^[29,30] and very recently following the sublimation of pure solid H₂CO₃ by means of matrix isolation FT-IR spectroscopy.^[31] Pure solid H₂CO₃ is accessible by five different techniques: (1) high-energy H or He irradiation of CO_2/H_2O ice mixtures,^[3,32-35] (2) high-energy UV irradiation of CO_2/H_2O ice mixtures,^[36] (3) H implantation in pure CO_2 ice,^[34,37] (4) the reaction of non-energetic OH radicals with CO^[38] and (5) acid-base chemistry in both (a) aqueous and (b) methanolic solution at cryo temperatures.^[5,8,14,39-43] As a surface-adsorbed species, carbonic acid was also identified on carbonate minerals in acidic but dry atmosphere^[44-46] and is a possible surface intermediate in the process of CO₂ adsorption on hydroxylated oxides, e.g. Fe₂O₃ or Al₂O₃.^[47,48]

The main analytical tool for the detection and characterization of solid carbonic acid has been infrared spectroscopy. Techniques (1), (2), (3)^[3,32–37] and (5a)^[5,8,40,42] (in aqueous environment) yield practically identical infrared spectra which are assigned to crystalline β -carbonic acid. By contrast, technique (5b) (in methanolic solution) results in a distinct spectrum assigned to crystalline α -carbonic acid.^[39,41,42] Other polymorphs of carbonic acid are not known at present, even though the infrared spectra obtained recently by Oba *et al*.^[38] might be interpreted to be different from the spectra of the α - and β -forms. The formation of these two crystalline polymorphs from solution is preceded by the formation of two distinct amorphous forms, in which the hydrogen-bond connectivity and conformational state are similar to the crystalline forms^[43] – reflecting Ostwald's step rule. The process of crystallization from the amorphous forms

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was also followed by powder X-ray diffraction.^[43] However, the quality of the powder diffractograms in the crystalline state is currently not sufficient to solve the crystal structures of the two polymorphs using techniques such as Rietveld refinement. Thus, 'no information on either the local or long-range structure' of the carbonic acid polymorphs is known as pointed out by Tossell.^[49]

Here, we report the Raman spectra of α - and β -carbonic acid in the range \sim 90–4000 cm⁻¹. This is a continuation of our recent Raman study on amorphous and β -carbonic acid, which aimed at the question whether amorphous and/or β -carbonic acid might be present on the icy parts of the Mars surface.^[14] Raman spectra of α carbonic acid are presented here for the first time, and the Raman spectra of β -carbonic acid are extended to lower wavenumbers. Raman and IR spectra yield complementary information, as IRforbidden transitions may be Raman active and vice versa. This is called 'The rule of mutual exclusion'^[50-54] and obeyed strictly in case of β -carbonic acid. By contrast, we here find that bands are both Raman and IR active in case of α -carbonic acid. This is suggestive of the centrosymmetric cyclic dimer as the local structural motif in β -carbonic acid and a non-centrosymmetric motif in α -carbonic acid. Based on the observed band positions and on the observation of needle-like crystals under the light microscope, we suggest a chain-like catemer motif or a sheet-like motif for α -carbonic acid. This finding is supported by normal mode analysis in dimeric carbonic acid units. Whereas we have focused on the 400-4000 cm⁻¹ region in our earlier study, the comparison with theory focuses on the low wavenumber region at <400 cm⁻¹ here. In addition, also a comparison with other carboxylic acids such as oxalic acid or malonic acid is presented.

Experimental

Solid H₂CO₃ was prepared via a cryogenic technique (5) by protonation of hydrogencarbonate anions. For details of the technique, see Refs [40,41], and for a schematic drawing of the apparatus see Ref. [55]. For the preparation of α -carbonic acid (technique 5b) methanolic solutions of 0.4 M KHCO₃ and 1.5 M HCI and for β -carbonic acid (technique 5a) aqueous solutions of 0.1 м КНСО₃ and 1 м HBr were used [КНСО₃ (Merck p. a.), CH₃OH (Aldrich, HPLC quality), HBr (Fluka, 48% p. a.), HCl (Supelco, 3 M in CH₃OH)]. Thin films ($<4 \,\mu$ m) of crystalline carbonic acid were prepared by alternately vitrifying droplets of the acidic and basic solutions on optical CsI windows (13 \times 2 mm) held at 77 K. By heating the films in vacuo ($\sim 10^{-7}$ mbar) to 200 K (methanolic solution) or 220 K (aqueous solution), the protonation of KHCO₃ takes place and the solvent and excess acid are removed from the window. To confirm the removal of the solvent and formation of crystalline carbonic acid, infrared spectra were recorded in situ in transmission. For this purpose, the beam passes through two KBr windows (49 × 6 mm) sealing the vacuum chamber and the sample on the CsI window inside the chamber. IR spectra were recorded on a Varian Excalibur FT-IR spectrometer at 4 cm⁻¹ resolution, by co-adding 100 scans.

After breaking the vacuum, the CsI window with the H_2CO_3 film was transferred under liquid nitrogen to a precooled (~80 K) microscope cryostat (Microstat N, Oxford) in order to carry out Raman measurements, for which we employ a Labram-1B 10/78IM instrument from Dilor S.A. The excitation wavelength of 632.8 nm is supplied by a 20 mW He–Ne laser. The laser light is totally reflected by a dichroic mirror (Super Notch Plus filter) through a quartz window of the cryostat towards the sample placed below an Olympus BX 40 microscope (using a LM Plan $50 \times$ NA 0.50 objective). The Raman scattered light is collected by the microscope and totally transmitted through the notch filter towards the confocal hole and the entrance slit of the spectrometer. Then a spectrum is formed on a CCD detector. A 1800 grooves mm⁻¹ grating with a resolution of $\sim 2 \text{ cm}^{-1}$ was used. The overall recording time for the spectra shown in Figs 1 and 2 was 7 min. The abscissa was calibrated with a silicon standard, and the sharp Raman shifts are accurate to $\pm 2 \text{ cm}^{-1}$ for the 1800 g mm⁻¹ grating. The relative intensities of the bands in different parts of the figures are not shown on the same scale. An Oxford Microstat was used as the cryostat. The temperature of the sample was controlled by a LakeShore Cl330 autotuning temperature controller and remained constant to within $\pm 0.2 \text{ K}$. For imaging of the surface, a TV camera is used.

Since the crystal structure is unknown, we do not attempt an assignment of the Raman bands to normal modes of a particular symmetry type. Therefore, the assignment can only be attempted by a qualitative description of the modes, similar to that of the IR spectra of H₂CO₃ polymorphs, ^[3,5,32,33,36,40,41,43,56] and it is based on comparisons with assignments of other crystalline carboxylic acids and on our calculation of the harmonic wavenumbers in carbonic acid oligomers. We note that strong coupling between the various modes is often observed for carboxylic acid oligomers, which makes vibrational assignments in terms of distinct modes even less meaningful.^[57-59] The vibrational spectra, including Raman and IR intensities, of carbonic acid oligomers were calculated using the Gaussian09 and Gaussian98 suite of programs. Carbonic acid dimers were calculated at a range of levels of theory, including MP2/aug-cc-pVDZ, B3LPY/6-31G(d), B3LYP/augcc-pVDZ and B3LYP/6-31+G(d). In this work, we pay particular attention to the Raman active modes in the range 100-500 cm⁻¹. In this range, variation of the basis set and/or method results in a shift of no more than ± 5 cm⁻¹. This good performance of the hybrid density functional and the low computational cost of this method allowed us to examine also the spectra of two possible tetramer configurations.

Results

Raman spectra of α -H₂CO₃ (Fig. 1, bottom; and Fig. 2, bottom) were recorded by focusing the laser on the crystal needles (Fig. 3, top). The infrared spectrum (Fig. 1, top) was recorded directly before transferring the sample window to the Raman microstat. The vertical dashed lines are shown as a guide to the eye where the intensity is seen both in the Raman and IR in Fig. 1. Compared to the Raman spectrum of β -H₂CO₃ (reported in Ref. [14]), more bands are detected in case of α -H₂CO₃ at 500–4000 cm⁻¹. This immediately suggests a lower symmetry of the latter. The assignment of Raman bands is presented in Table 1 and is based on the assignment of IR bands presented in Ref. [41] because of the appearance of the bands in the Raman spectrum at very similar positions as in the infrared spectrum. The assignment of IR bands in Ref. [41] is itself based on IR spectra of different carbonic acid isotopologues. In addition to this assignment, we also refer to the assignment of the acid I to acid VII fundamentals presented by Wolfs and Desseyn on the example of some dicarboxylic acids.^[59] We wish to stress that the assignments made in Table 1 do not imply that we are dealing with uncoupled modes. Instead, most modes are coupled with other modes, as also found in other (di)carboxylic acids such as malonic acid,^[57] and so the assignments represent only qualitative descriptions.



Figure 1. Comparison between FT-IR spectrum (top) and Raman spectrum (bottom) of α -H₂CO₃. The comparison between FT-IR spectrum and Raman spectrum of β -H₂CO₃ is published as Fig. 2 in Kohl *et al., Angew. Chem. Int. Ed.* **2009**, *48*, 2690.

ν (cm ⁻¹)	Raman intensity	Assignment	Di- and polycarboxylic acids, Raman at 20 °C, Refs [59,67]
3057, 3039	m, m	ν(O-H)	
2976	S	ν(O-H)	
1630	vw	ν (C=O), acid I	1647–1782
1489, 1465, 1447, 1427	m, m, w, vw	$v_{as}[C(OH)_2]$, acid II	1367–1437
1203	m	δ (COH), acid III	1219–1275
1164	W		
1093	W	$v_s[C(OH)_2]$	
928	VS	π (OH), δ_{oop} (COH· · · O), acid IV	915–950 (IR)
799	W	$\delta_{oop}(CO_3)$	
681	m	δ (COO), acid V	641–684
567, 558	S, S	ρ (COO), δ_{ip} (CO ₃), acid VI	442-601
149			
134			
113			

In the OH stretching region, we find two intense Raman bands at 3057 and at 3039 cm⁻¹ as well as a very strong band at 2976 cm⁻¹, which we detect as weak bands in the IR. By contrast, we cannot find the IR overtones at 2701 and 2593 cm⁻¹ in the Raman spectrum.^[41] The absence of overtones in the Raman spectrum was expected because only a very small fraction of photons is scattered inelastically in the Raman process. The absence of these bands, therefore, corroborates the earlier assignment of these IR bands to overtones and combination bands (see Table 1 in Ref. [41]). We attribute the Raman band centered in Fig. 1 at \sim 3090 cm⁻¹ to the O–H stretching vibration of hexagonal ice (ice *Ih*)^[60] that

had condensed onto the film during the transfer of the film and the disk from the IR apparatus to the Raman microstat. Also, the translational mode of ice *Ih* centered at 225 cm⁻¹ is detected in this film (Fig. 2, bottom).^[61,62] The band at ~3090 cm⁻¹ is also observable in the Raman spectra of D₂CO₃ samples (not shown) and shows a significantly higher full width at half-maximum, which is consistent with our assignment of this band to ice *Ih*. We note that the carboxylic acid OH stretching bands are generally weak in the Raman spectrum.^[58] Whereas we were unable to detect it reliably in case of β -H₂CO₃,^[14] we did detect (at least) three sharp bands in case of α -H₂CO₃. We interpret these bands due

RAMAN SPECTROSCOPY



Figure 2. Low wavenumber Raman modes (<500 cm⁻¹) in β -H₂CO₃ (top) and α -H₂CO₃ (bottom).



Figure 3. Microscopy images of α -H₂CO₃ (top) and β -H₂CO₃ (bottom). The images were recorded at 80 K in the Raman microstat using the confocal Raman microscope equipped with a 50× ULWD objective. Raman spectra of α -carbonic acid as depicted in Figs 1 and 2 were obtained only when focusing directly on the needles, whereas no bands at all were obtained when focusing slightly beneath the needles. Raman spectra of β -carbonic acid as shown in Fig. 2 and in Ref. [14] were obtained by focusing the laser on the dark sites.

to the OH groups in (at least) three different environments, two of which might be similar because of the very small wavenumber shift and one being inherently different. Tentatively, we assign two weakly intermolecular-bonded OH groups and one more strongly hydrogen-bonded OH group at lower wavenumbers. The picture of two weakly interacting catemer chains appears. Every chain has the same kind of intrachain hydrogen bonds and OH groups sticking out alternatingly to the left and right of the chain direction. These OH groups might loosely interact with the neighboring chains, therebye producing sheets of carbonic acid.

The intense and structured carbonyl stretching band in the IR spectrum of α -H₂CO₃ is centered at 1711 cm⁻¹. The C=O stretching mode is observed typically in the 1780 cm⁻¹ region for isolated carboxylic acids and is shifted to 1720 cm⁻¹ on hydrogenbond formation.^[59] In the Raman spectrum, the C=O stretch is

typically observed at much lower intensity in most carboxylic acid polymorphs^[59] and shifted to lower wavenumbers. The difference, which is also referred to as factor-group splitting, is typically on the order of $30-50 \text{ cm}^{-1}$. In case of carbonic acid, we observe a very weak Raman band at \sim 1630 cm⁻¹, which implies a splitting $\Delta \nu$ (C=O) of ~80 cm⁻¹. The large splitting between the Raman v_s C=O mode and the IR v_{as} C=O mode is characteristic for carboxylic acid dimers^[59] and has been examined theoretically for the formic acid dimer.^[63,64] The $\Delta \nu$ (C=O) value ~80 cm⁻¹ is consistent with splittings obtained in various calculations of the carbonic acid dimer.^[14,65] If the carbonyl group in the acid were not hydrogen-bonded or formed only (an) internal hydrogen bond(s), then one would expect a band in the range 1705-1760 cm⁻¹ in both the IR and Raman spectra.^[66,67] The absence of the Raman band in this region and the large splitting are strong indications for the presence of a dimer or a longer polymeric chain in α -H₂CO₃. Different polymorphs of di- and polycarboxylic acids such as oxalic acid, malonic acid or citric acid show the Raman active antisymmetric stretch ν (C=O) at 1647–1782 cm⁻¹.^[59,67] As a consequence of the OH group directly connected to the carboxyl group rather than a carbon atom, both α -H₂CO₃ (1630 cm⁻¹) and β -H₂CO₃ (1608 cm⁻¹) show weak Raman bands at even lower wavenumbers. We, therefore, assign this band as the acid I band in Table 1 and suggest a chainlike catemer motif for α -H₂CO₃.

The simultaneous appearance of bands in the Raman and IR is best evidenced at $500-1500 \text{ cm}^{-1}$. All bands observed in the Raman spectrum in this region can also be observed at very similar positions in the IR. These are marked by vertical dashed lines in Fig. 1. Bandsplitting is apparent in both IR and Raman spectra, especially in the COH deformation region between 1430 and 1490 cm⁻¹, where four bands separated by approximately 20 cm⁻¹ are observed in both spectra. Also, this band system falls outside the range typically seen in other carboxylic or dicarboxylic acids, namely 1367 - 1437 cm⁻¹ (cf Table 1). This might again be explained in terms of the second OH group directly attached to the C=O group, and so we assign these bands as the acid II bands. Bands at 1203 and 1164 cm⁻¹ are intense in Raman but weak in the IR. The intense IR band at 1087 cm⁻¹ has a Raman counterpart at 1093 cm⁻¹. The most intense Raman band at 928 cm⁻¹ can be attributed to the split band in the IR spectrum at 912/920 cm⁻¹. The IR active band at 803 cm⁻¹ has its counterpart at 799 cm⁻¹ in the Raman. The Raman bands in the skeletal region at 681 cm⁻¹ and the doublet at 567 and 558 cm⁻¹ are shifted by ca. $+20 \text{ cm}^1$ to 702 and 584 cm⁻¹. The latter band is only resolved as a doublet in the Raman but not in the IR. All these coincidences between Raman and IR clearly show that the principle of mutual exclusion does not hold in case of α - H_2CO_3 , in contrast to the situation in β - H_2CO_3 .^[14] A local center of inversion and the cyclic centrosymmetric dimer of carbonic acid (depicted in Fig. 4(a)) can clearly be ruled out to be the building block of α -H₂CO₃, whereas we suggest this dimer to be the building block of β -H₂CO₃.^[14] The large factor-group splitting in case of α -H₂CO₃ suggests that other types of dimers, oligomers or chainlike structures rather than well-separated monomers build α -H₂CO₃. We also want to emphasize that we have interpreted the IR spectra to indicate the presence of two different conformational or H-bonded states of carbonic acid.^[41] This assessment has been based on the observation of two intense IR band systems assignable to the asymmetric COH stretching modes, namely at 1311/1327 cm⁻¹ and 1422/1447/1465/1479 cm⁻¹, respectively (Fig. 1, top). In the Raman spectra, only the latter band system,



Figure 4. (a) Local structural motif suggested for β -H₂CO₃ containing a center of inversion (C_{2h} point group). (b) Candidate local structural motif suggested for α -H₂CO₃ without a center of inversion (C_{S} point group symmetry).

but not the former, is observed. We explain this in terms of at least two carbonic acid entities in the primitive unit cell, and we assume that this is due to an element of symmetry in the unit cell other than a center of inversion. If we follow the picture of two catemer chains connected by weak interchain hydrogen bonds, then there are two different $C(OH)_2$ stretch vibrations in every chain: one with two hydrogen-bonded OH groups, and the other with one hydrogen-bonded OH group. We expect the completely interconnected $C(OH)_2$ group at lower wavenumbers and an increase of the full width at half-maximum due to the interaction, as observed in the FT-IR spectra. At the same time, we expect a stronger change of the symmetry for these groups. This might be the reason for its disappearance in the Raman spectra, in accordance with the symmetry selection rules.

Murillo et al.^[68] have suggested a large number of different dimers, which might serve as the fundamental dimeric unit building such a catemer chain. We have selected among these some energetically rather low-lying dimers, which are capable of producing chainlike structures and which do not show a center of inversion. One candidate structure, which we regard as a good candidate building block of α -H₂CO₃ is depicted in Fig. 4(b). We wish to emphasize, though, that also other dimers not containing a center of inversion are possible as building blocks of α -H₂CO₃. For instance, we also regard a crystal structure built from *cis-trans* carbonic acid monomers as consistent with the Raman spectroscopic results. These monomers could build linear chains via single (stronger) hydrogen bonds connecting one OH group from the first carbonic acid molecule with the C=O group of the second carbonic acid molecule. The second OH group in each carbonic acid molecule would link one linear chain of carbonic acid molecules with the other by forming (weaker) hydrogen bonds.

Recently, some candidate crystal structures have been predicted by Reddy *et al.*^[69] and by us.^[43] From the considerations mentioned above, the crystal structures shown in Figs. 1 and 5 in Ref. [69] and

in Fig. 5(b) and (c) in Ref. [43] are possible candidate structures for β -H₂CO₃ but not for α -H₂CO₃. By contrast, the crystal structures in Figs. 2–5 in Ref. [69] and in Fig. 5(a) and (d) in Ref. [43] are possible candidate structures for α -H₂CO₃ but not for β -H₂CO₃. In the absence of IR and Raman solid state spectra calculated from various candidate structures, the experimental IR and Raman data do not allow us to pick a specific crystal structure as the best candidate at present. However, they allow restricting which kind of structures are possible and which are not possible from symmetry considerations.

Low wavenumber Raman spectra

It is well known that low wavenumber Raman spectra show highly characteristic patterns allowing easy discrimination between different crystals. This fingerprint information was recently employed to clearly distinguish between different solid nitric acid hydrates, including metastable polymorphs.^[70,71] At <400 cm⁻¹, β -H₂CO₃ shows more Raman bands than α -H₂CO₃ (see Fig. 2). This is in contrast to the situation encountered at $>400 \text{ cm}^{-1}$ (see discussion above). While we explain the higher number of Raman bands at $>400 \text{ cm}^{-1}$ in α -H₂CO₃ in terms of missing local symmetry, the low number of Raman bands at <400 cm⁻¹ hints at a rather simple unit cell for α -H₂CO₃. By contrast, the unit cell in β -H₂CO₃ could be more complex and contain more carbonic acid molecules. The lower limit for internal modes in a dicarboxylic acid is around 250 cm⁻¹.^[57] Whereas there are bands at 250–500 cm⁻¹ in case of β -H₂CO₃,^[14] there is none in case of α -H₂CO₃. There are only four bands at 225, 149, 134 and 113 cm⁻¹ detected at $<400 \text{ cm}^{-1}$ in α -H₂CO₃ (see Table 3). As mentioned above, the band at 225 cm⁻¹ (Fig. 2, bottom) can also arise from the translational band in condensed hexagonal ice. By contrast, at least ten bands appear in case of β -H₂CO₃ in this region (see Fig. 2, top and Table 2). This includes the most intense Raman band appearing at 192 cm⁻¹, which is missing in α -H₂CO₃. A Raman band at 194 cm⁻¹ was recently detected in supersonic jet expansion studies of formic acid and attributed to the intermonomer stretching mode in cyclic, centrosymmetric dimers of formic acid, i.e. the RC(OHO)₂CR entity with R=H.^[72] Also, in matrix isolation studies of gas-phase formic acid, an A_a band was found at 196 cm⁻¹, which was attributed to this mode.^[73] The cyclic centrosymmetric dimer of carbonic acid contains the same RC(OHO)₂CR subunit with R=OH, and so this band might be attributed to the inter-monomer stretch in the cyclic dimer of C_{2h} symmetry (see Fig. 4(a)). Indeed, we calculate the Raman active inter-monomer normal mode of A_a type to be 190 \pm 2 cm⁻¹ at B3LYP and MP2 levels of theory using 6-31 + G(d) or aug-ccpVDZ basis sets from the geometry shown in Fig. 4(a). We thus suggest the mode at 192 cm⁻¹ observed in β -H₂CO₃ (Fig. 2, top) to be the inter-monomer stretching mode. Also, the intense modes observed at 165 and 125 cm⁻¹ agree very well with the predicted Raman modes at 166 \pm 4 and 125 \pm 5 cm⁻¹ (see Table 2). By contrast, the modes at 176, 144 or 110 cm⁻¹ are not predicted from this simplified gas-phase calculation and might necessitate consideration of larger clusters, inclusion of crystal symmetry and a proper description of the crystal field. The mode at 144 cm^{-1} , for instance, appears in our calculation of the centrosymmetric carbonic acid tetramer (see Table 2). Calculation of vibrational spectra for all at least 200 different candidate crystal structures^[43] are, however, not feasible.

The harmonic frequency calculations on the dimer depicted in Fig. 4(b) show two Raman active bands in the range ?

?

110

92



Table 2.Lowobtained at ~ 80	wavenumber) K	Raman modes (cm ⁻¹) of β -H ₂ CO ₃
Experimental v (cm ⁻¹)	Raman intensity	Calculated (Fig. 4(a))	β -Oxalic acid Ref. [76]
258	m, br		
192	VS	190 \pm 2 (A $_g$)	189 (A _g)
176	S		167/187 (B _g)
165	sh	$166\pm4~(A_g)$	159 (A _g)
144	S		
138	sh		
125	?	$121 \pm 2 (B_q)$	131 (<i>B</i> _q)

Note that experimental Raman intensities decay considerably at <150 cm⁻¹ because of the use of a strong filter for the intense Rayleigh scattering (at 0 cm⁻¹). Calculated wavenumbers refer to gas-phase calculations in the harmonic approximation for the geometry depicted in Fig. 4(a). Error bars are given from a comparison of B3LYP and MP2 calculations and a comparison of different basis sets. Data at 300 K are from Ref. [76] and are corrected by $+10 \text{ cm}^{-1}$ in order to correct for the temperature change from 300 K to ~80 K.

Table 3. Low wavenumber Raman modes (cm $^{-1}$) of $\alpha\text{-H}_2\text{CO}_3$ obtained at ${\sim}80~\text{K}$					
Experimental ν (cm ⁻¹)	Raman intensity	Calculated (Fig. 4(b))	α-Oxalic acid Ref. [76]		
225	m, br	220 (A')			
149	VS		161		
134	S	135 (A')	127/138		
113	m		117		

Note that experimental Raman intensities decay considerably at $<150 \text{ cm}^{-1}$ because of the use of a strong filter for the intense Rayleigh scattering (at 0 cm⁻¹). Calculated wavenumbers refer to gas-phase calculations in the harmonic approximation for the geometry. Most intense bands from Ref. [76] are given (90 K data).

 $100-500 \text{ cm}^{-1}$, namely two A' modes at 135 and 220 cm⁻¹. Most noteworthy, the mode at 192 cm⁻¹ is clearly absent in such type of dimer and, furthermore, the two predicted Raman bands can also be found in the measured spectrum of α -H₂CO₃ (Fig. 2, bottom and Table 3). We wish to emphasize that the observed mode at 225 cm⁻¹ might also arise from hexagonal ice condensed during sample transfer (see above). However, the broad, tailed nature of this band is also consistent with the presence of both hexagonal ice and carbonic acid dimer translational bands. In fact, band deconvolution is consistent with two weakly separated bands producing the band shape in Fig. 2, bottom. Of course, we have also considered other types of dimers, trimers and tetramers including the non-planar T-shaped dimer similar to the structures depicted in Fig. 4 in Ref. [68]. However, none of these can explain the vibrational features in the low wavenumber Raman spectrum of β -H₂CO₃ and α -H₂CO₃ better than the structures depicted in Fig. 4(a) and (b), respectively.

The main finding of the absence of the Raman band at 192 cm⁻¹ in crystal structures based on non-centrosymmetric dimers is also found in the related case of oxalic acid. Just like in carbonic acid, there are also two polymorphs in oxalic acid.^[74] And just like in carbonic acid, the crystal structure of the β -polymorph is based on

the centrosymmetric dimer, whereas the crystal structure of the α -polymorph is not based on the centrosymmetric motif (see Fig. 1 in Ref. [75]). The Raman spectra of both polymorphs, including the low wavenumber region, were measured by DeVillepin and coworkers.^[76] The most intense Raman modes in α -oxalic acid were found at 148, 130, 121 and 110 cm^{-1} at 300 K, which shifted by +7 to +13 cm⁻¹ when cooling the sample to 90 K. This compares very well with the Raman spectrum of α -carbonic acid (Fig. 2, bottom). In case of β -oxalic acid, two A_q modes are observed at 179 and 149 cm⁻¹ at 300 K. The data at 90 K are not reported, but assuming a shift of +10 cm⁻¹ these two modes would be at 189 and 159 cm⁻¹, in excellent agreement with the observed spectrum of β -carbonic acid (Fig. 2, top). Also the three B_q modes observed in β -oxalic acid at 121, 157 and 177 cm⁻¹ are quite similar to the modes observed in β -carbonic acid (Fig. 2, top), most notably the B_q mode at 121 cm⁻¹, which also appears in the gas-phase calculation of the cyclic dimer. We, thus, conclude that the observed low wavenumber Raman spectra are also consistent with the suggestion of the centrosymmetric dimer (Fig. 4(a)) as the local motif in β -H₂CO₃ and a non-centrosymmetric dimer as the local motif in α -H₂CO₃ (Fig. 4(b)). Whether this local motif results in a sheet-like crystal structure or a chain-like catemer type one cannot be answered from the Raman data reported here and needs to be clarified in future work.

Discussion/Conclusions

In summary, we report here the Raman spectra of α -H₂CO₃ including the low wavenumber range. In addition, we report new low wavenumber Raman data for β -H₂CO₃. While this type of Raman spectra has been reported between 30 and 100 years ago for other carboxylic or dicarboxylic acids such as formic acid,^[77,78], oxalic acid,^[76,79-81] malonic acid^[57,79,82] or acetic acid,^[67] the smallest organic acid, containing only one C atom, has eluded attempts of Raman characterization until 2009.^[14] Because of the possible astrophysical relevance of carbonic acid^[2-15] and the possibility of exploring astrophysical objects such as Mars using Raman spectroscopy,^[83-88] the data are of importance in possibly identifying α -H₂CO₃ in outer space. For this purpose, the most intense band at 928 cm⁻¹ seems to be the best marker band. For comparison, the most intense band in β -H₂CO₃ is located at 1054 cm⁻¹, while there is no band close to 928 cm⁻¹. Because of the separation of >100 cm⁻¹ and the high intensity of the two bands, the α - and β -polymorphs of carbonic acid can be easily discriminated by Raman spectroscopy. For comparison, the most intense Raman band in aqueous carbonic acid attributed to the C-OH stretching vibration is at 1017 cm^{-1} .^[26]

The interpretation of vibrational spectra provides valuable data, in particular concerning the local short-range ordering in these polymorphs. A comparison between IR and Raman spectra reveals that the principle of mutual exclusion only holds for β -H₂CO₃ but not for α -H₂CO₃. This suggests a local center of inversion in the former and the absence of such a center in the latter. Comparison of the observed Raman spectra with gas-phase calculations of simple carbonic acid dimers (*cf.* Fig. 4) is consistent with the idea of the centrosymmetric dimer shown in Fig. 4(a) building the β -polymorph and a non-centrosymmetric dimer similar to the one shown in Fig. 4(b) building the α -polymorph. A similar situation of the cyclic, centrosymmetric RC(OHO)₂CR motif is encountered in β -oxalic acid (R=COOH),^[76] and interestingly its low wavenumber Raman pattern observed resembles the one observed here for β -carbonic acid (R=OH). Most notably, the inter-monomer stretching mode is observed in both cases at approximately 190 cm⁻¹. Similarly, the inter-monomer stretching mode in gas-phase formic acid (R=H) is also observed at 194 cm⁻¹, thus confirming the assignment of the intense Raman mode in β -H₂CO₃ at 192 cm⁻¹.^[72] By contrast, this mode is missing in both α -oxalic acid and α -carbonic acid. Whereas it is known that the former has a sheet-like crystal structure built from oxalic acid dimers connected by hydrogen bonds at heads and tails, we infer from the present data that a similar situation is encountered in case of α -carbonic acid. Whether α -carbonic acid can be characterized on the long range by chain-like structures or by planar or wavy sheets and which monomer conformation is present cannot be answered from the present work. All of these possibilities can be found among the predicted candidate structures for carbonic acid,^[43,69] and elucidation of the crystal structure will require future work. The Raman spectra presented here constitute an important guide for the process of determining the crystal structures of the two known polymorphs of carbonic acid, which is probably one of the last fundamental small molecules for which the crystal structure is unknown.

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References

- [1] T. Loerting, J. Bernard, ChemPhysChem 2010, 11, 2305.
- [2] J. S. Lewis, D. H. Grinspoon, Science 1990, 249, 1273.
- [3] N. DelloRusso, R. K. Khanna, M. H. Moore, J. Geophys. Res. Planets 1993, 98, 5505.
- [4] R. K. Khanna, J. A. Tossell, K. Fox, *lcarus* 1994, 112, 541.
- [5] W. Hage, A. Hallbrucker, E. Mayer, J. Chem. Soc., Farad. Trans. 1995, 91, 2823.
- [6] J. R. Brucato, A. C. Castorina, M. E. Palumbo, M. A. Satorre, G. Strazzulla, Planet. Space Sci. 1997, 45, 835.
- [7] G. Strazzulla, J. R. Brucato, G. Cimino, M. E. Palumbo, *Planet. Space Sci.* **1996**, 44, 1447.
- [8] W. Hage, K. R. Liedl, A. Hallbrucker, E. Mayer, Science 1998, 279, 1332.
- [9] M. L. Delitsky, A. L. Lane, J. Geophys. Res. [Planets] 1998, 103, 31391.
- [10] R. L. Hudson, M. H. Moore, J. Geophys. Res. [Planets] 2001, 106, 33275.
- [11] M. H. Moore, R. L. Hudson, P. A. Gerakines, *Spectrochim. Acta* **2001**, *57A*, 843.
- [12] M. H. Moore, R. L. Hudson, R. F. Ferrante, *Earth, Moon, and Planets* 2003, 92, 291.
- [13] R. L. Hudson, J. Chem. Educ. 2006, 83, 1611.
- [14] I. Kohl, K. Winkel, M. Bauer, K. R. Liedl, T. Loerting, E. Mayer, Angew. Chem., Int. Ed. 2009, 48, 2690.
- [15] Z. Peeters, R. L. Hudson, M. H. Moore, A. Lewis, *lcarus* 2010, 210, 480.
- [16] A. Khanna, N. A. Kurtzman, J. Nephrol. 2006, 19, S86.
- [17] P. Swietach, R. D. Vaughan-Jones, A. L. Harris, *Cancer Metastasis Rev.* 2007, 26, 299.
- [18] R. D. Vaughan-Jones, K. W. Spitzer, P. Swietach, J. Mol. Cell. Cardiol. 2009, 46, 318.
- [19] G. Liamis, H. J. Milionis, M. Elisaf, Drug Saf. 2010, 33, 371.
- [20] C. L. Sabine, R. A. Feely, N. Gruber, R. M. Key, K. Lee, J. L. Bullister, R. Wanninkhof, C. S. Wong, D. W. R. Wallace, B. Tilbrook, F. J. Millero, T.-H. Peng, A. Kozyr, T. Ono, A. F. Rios, *Science* **2004**, *305*, 367.
- [21] R. A. Feely, C. L. Sabine, K. Lee, W. Berelson, J. Kleypas, V. J. Fabry, F. J. Millero, *Science* **2004**, *305*, 362.
- [22] J. C. Orr, V. J. Fabry, O. Aumont, L. Bopp, S. C. Doney, R. A. Feely, A. Gnanadesikan, N. Gruber, A. Ishida, F. Joos, R. M. Key, K. Lindsay, E. Maier-Reimer, R. Matear, P. Monfray, A. Mouchet, R. G. Najjar, G.-K. Plattner, K. B. Rodgers, C. L. Sabine, J. L. Sarmiento, R. Schlitzer,

R. D. Slater, I. J. Totterdell, M.-F. Weirig, Y. Yamanaka, A. Yool, *Nature* **2005**, *437*, 681.

- [23] P. Glynn, *Rev. Mineral. Geochem.* **2000**, *40*, 481.
- [24] J. W. Morse, R. S. Arvidson, Earth-Sci. Rev. 2002, 58, 51
- [25] R. Kruse, E. U. Franck, Ber. Bunsenges. Phys. Chem. 1982, 86, 1036.
- [26] H. Falcke, S. H. Eberle, Water Res. 1990, 24, 685.
- [27] K. Adamczyk, M. Premont-Schwarz, D. Pines, E. Pines, E. T.J. Nibbering, Science 2009, 326, 1690.
- [28] J. K. Terlouw, C. B. Lebrilla, H. Schwarz, Angew. Chem., Int. Ed 1987, 26, 354.
- [29] T. Mori, K. Suma, Y. Sumiyoshi, Y. Endo, J. Chem. Phys. 2009, 130, 204308.
- [30] T. Mori, K. Suma, Y. Sumiyoshi, Y. Endo, J. Chem. Phys. 2011, 134, 044319.
- [31] J. Bernard, M. Seidl, I. Kohl, E. Mayer, K. R. Liedl, O. Galvez, H. Grothe, T. Loerting, Angew. Chem. Int. Ed. 2011, 50, 1939.
- [32] M. H. Moore, R. Khanna, B. Donn, J. Geophys. Res. [Planets] 1991, 96, 17541.
- [33] M. H. Moore, R. K. Khanna, Spectrochim. Acta 1991, 47A, 255
- [34] J. R. Brucato, M. E. Palumbo, G. Strazzulla, Icarus 1997, 125, 135.
- [35] G. Strazzulla, G. Leto, F. Spinella, O. Gomis, *Astrobiology* **2005**, *5*, 612.
- [36] P. A. Gerakines, M. H. Moore, R. L. Hudson, Astron. Astrophys. 2000, 357, 793.
- [37] M. Garozzo, D. Fulvio, O. Gomis, M. E. Palumbo, G. Strazzulla, *Planet. Space Sci.* 2008, *56*, 1300.
- [38] Y. Oba, N. Watanabe, A. Kouchi, T. Hama, V. Pirronello, Astrophys. J. 2010, 722, 1598.
- [39] W. Hage, A. Hallbrucker, E. Mayer, J. Am. Chem. Soc. 1993, 115, 8427.
- [40] W. Hage, A. Hallbrucker, E. Mayer, J. Chem. Soc., Farad. Trans. 1996, 92, 3197.
- [41] W. Hage, A. Hallbrucker, E. Mayer, J. Chem. Soc., Farad. Trans. 1996, 92, 3183.
- [42] W. Hage, A. Hallbrucker, E. Mayer, J. Mol. Struct. 1997, 408, 527.
- [43] K. Winkel, W. Hage, T. Loerting, S. L. Price, E. Mayer, J. Am. Chem. Soc. 2007, 129, 13863.
- [44] H. A. Al-Hosney, V. H. Grassian, J. Am. Chem. Soc. 2004, 126, 8068.
- [45] H. A. Al-Hosney, S. Carlos-Cuellar, J. Baltrusaitis, V. H. Grassian, Phys. Chem. Chem. Phys. 2005, 7, 3587.
- [46] H. A. Al-Hosney, V. H. Grassian, Phys. Chem. Chem. Phys. 2005, 7, 1266.
- [47] J. Baltrusaitis, J. D. Schuttlefield, E. Zeitler, J. H. Jensen, V. H. Grassian, J. Phys. Chem. C 2007, 111, 14870.
- [48] J. Baltrusaitis, V. H. Grassian, J. Phys. Chem. A 2010, 114, 2350.
- [49] J. A. Tossell, Inorg. Chem. 2006, 45, 5961.
- [50] T. Venkatarayudu, J. Chem. Phys. 1954, 22, 1269.
- [51] A. R. H. Cole, J. R. Durig, J. Raman Spectrosc. 1975, 4, 31.
- [52] R. Mendelsohn, E. U. Monse, J. Chem. Educ. 1981, 58, 582.
- [53] S. F. Parker, H. Herman, Spectrochim. Acta 1997, 53A, 119.
- [54] R. L. Frost, M. L. Weier, J. Raman Spectrosc. 2003, 34, 776.
- [55] I. Kohl, L. Bachmann, A. Hallbrucker, E. Mayer, T. Loerting, *Phys. Chem. Chem. Phys.* **2005**, *7*, 3210.
- [56] G. Strazzulla, G. A. Barratta, M. E. Palumbo, M. A. Satorre, Nucl. Instrum. Methods Phys.Research, Sect. B 2000, 166–167, 13.
- [57] D. Bougeard, J. De Villepin, A. Novak, Spectrochim. Acta 1988, 44A, 1281.
- [58] D. Lin-Vien, N. B. Colthup, W. G. Fateley, J. G. Grasselli, *The Handbook of Infrared and Raman Characteristic Frequencies of Organic Molecules*, Academic Press: New York, **1991**.
- [59] I. Wolfs, H. O. Desseyn, Appl. Spectrosc. 1996, 50, 1000.
- [60] P. T. T. Wong, E. Whalley, J. Chem. Phys. 1975, 62, 2418.
- [61] M. J. Taylor, E. Whalley, J. Chem. Phys. 1964, 40, 1660.
- [62] P. T. T. Wong, E. Whalley, J. Chem. Phys. 1976, 65, 829.
- [63] W. L. Qian, S. Krimm, J. Phys. Chem. **1996**, 100, 14602.
- [64] W. L. Qian, S. Krimm, J. Phys. Chem. A 2002, 106, 11663.
- [65] P. Ballone, B. Montanari, R. O. Jones, J. Chem. Phys. 2000, 112, 6571.
- [66] N. B. Colthup, L. H. Daly, S. E. Wiberley, Introduction to Infrared and Raman Spectroscopy, Academic Press: New York, 1964.
- [67] F. R. Dollish, W. G. Fateley, F. F. Bentley, Characteristic Raman Frequencies of Organic Compounds, Wiley: New York, 1974, pp 105.
- [68] J. Murillo, J. David, A. Restrepo, Phys. Chem. Chem. Phys. 2010, 12, 10963.
- [69] S. K. Reddy, C. H. Kulkarni, S. Balasubramanian, J. Chem. Phys. 2011, 134, 124511.
- [70] H. Grothe, C. E. L. Myhre, C. J. Nielsen, J. Phys. Chem. A 2006, 110, 171.



- [71] R. M. Escribano, D. Fernandez-Torre, V. J. Herrero, B. Martin-Llorente, B. Mate, I. K. Ortega, H. Grothe, *Vib. Spectrosc.* 2007, 43, 254.
- [72] P. Zielke, M. A. Suhm, Phys. Chem. Chem. Phys. 2007, 9, 4528.
- [73] A. Olbert-Majkut, J. Ahokas, J. Lundell, M. Pettersson, Chem. Phys. Lett. 2009, 468, 176.
- [74] L. J. Bellamy, R. J. Pace, Spectrochim. Acta 1963, 19, 435.
- [75] I. Nobeli, S. L. Price, J. Phys. Chem. A **1999**, 103, 6448.
- [76] J. De Villepin, A. Novak, D. Bougeard, *Chem. Phys.* **1982**, *73*, 291.
- [77] S. M. Blumenfeld, H. Fast, *Spectrochim. Acta* **1968**, *24*, 1449.
- [78] H. R. Zelsmann, Y. Marechal, A. Chosson, P. Faure, J. Mol. Struct. 1975, 29, 357.
- [79] D. Hadzi, N. Sheppard, Proc. Roy. Soc. (London) 1953, A216, 247.
- [80] J. De Villepin, A. Novak, Spectrochim. Acta 1978, 34A, 1019.

- [81] J. De Villepin, A. Novak, Spectrochim. Acta 1978, 34A, 1009.
- [82] J. De Villepin, M. H. Limage, A. Novak, N. Toupry, M. Le Postollec, H. Poulet, S. Ganguly, C. N. R. Rao, J. Raman Spectrosc. 1984, 15, 41.
- [83] A. Wang, L. A. Haskin, A. L. Lane, T. J. Wdowiak, S. W. Squyres, R. J. Wilson, L. E. Hovland, K. S. Manatt, N. Raouf, C. D. Smith, *J. Geophys. Res.*, [*Planets*] **2003**, *108*, 5005.
- [84] A. Ellery, D. Wynn-Williams, Astrobiology 2003, 3, 565.
- [85] J. L. Bishop, E. Murad, J. Raman Spectrosc. 2004, 35, 480.
- [86] J. Popp, M. Schmitt, N. Tarcea, W. Kiefer, M. Hilchenbach, T. Stuffler, S. Hofer, R. Hochleitner, A. Wuttig, R. Riesenberg, *Eur. Space Agency*, [Spec. Publ.] 2005, SP-588, 385.
- [87] P. Sobron, F. Sobron, A. Sanz, F. Rull, Appl. Spectrosc. 2008, 62, 364.
- [88] J. Jehlicka, A. Culka, H. G. M. Edwards, Planet. Space Sci. 2010, 58, 875.